

1 Stochastic similarities between the microscale of turbulence and 2 hydrometeorological processes

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7 Abstract

8 Turbulence is considered to generate and drive most geophysical processes. The simplest case is the
9 isotropic turbulence. In this paper, the most common three-dimensional power-spectrum-based
10 models of isotropic turbulence are studied in terms of their stochastic properties. Such models often
11 have a high-order of complexity, lack in stochastic interpretation and violate basic stochastic
12 asymptotic properties, such as the theoretical limits of the Hurst coefficient, in case that Hurst-
13 Kolmogorov behaviour is observed. A simpler and robust model (which incorporates self-similarity
14 structures, e.g. fractal dimension and Hurst coefficient) is proposed using a climacogram-based
15 stochastic framework and tested over high resolution observational data of laboratory scale as well as
16 hydrometeorological observations of wind speed and precipitation intensities. Expressions of other
17 stochastic tools like the autocovariance and power spectrum are also produced from the model and
18 show agreement with data. Finally, uncertainty, discretization and bias related errors are estimated for
19 each stochastic tool, showing lower errors for the climacogram-based ones and larger for power-
20 spectrum ones.

21 **Keywords:** isotropic-stationary turbulence; hydrometeorological processes; stochastic modelling;
22 climacogram; power spectrum; uncertainty-bias

23 1. Introduction

24 Turbulence originates from the Greek word ‘τύρβη’ (cf. ‘...τὴν τύρβην ἐν ἧ ζῶμεν’:‘...for the
25 turbulence in which we live’, Isokrates, 15.130) which means disorder, confusion, turmoil. Turbulence
26 is considered to generate and drive most geophysical processes, e.g. wind turbulence giving birth and
27 spatiotemporal variability in cloud rainfall (cf. Falkovich et al. 2002), yet it is regarded as mystery
28 within classical physics (McDonough 2007 ch. 1). Studying turbulent phenomena is of high
29 importance for hydrology (e.g. Mandelbrot and Wallis 1968, Rinaldo 2006) as the microscopic
30 processes (related to turbulence) can help understand the macroscopic ones (related to hydrology),
31 since they enable the recording of very long time-series and with a high resolution, a rare case for
32 hydrological processes (cf. Koutsoyiannis 2014). The simplest case of turbulent state (in terms of
33 mathematical calculations) is the stationary, isotropic and homogeneous turbulence. While this is a
34 physical phenomenon that has been recognized hundreds of years ago, still there is no universally
35 agreed mathematical definition for the so-called ‘turbulent state’ (Tessarotto and Ascì 2010). Leonardo
36 da Vinci tried to give a definition 500 years ago, based on his observations that water falling into a
37 sink forms large eddies as well as rotational motion (cf. Richter 1939). Interestingly, Heisenberg (1948)
38 commented on the definition of turbulent state of flow that it is just the result of infinite degrees of
39 freedom developed in a liquid flowing without friction and thus, by contrast, laminar flow is a state of
40 flow with reduced degrees of freedom caused by the viscous action. In 1880, Reynolds introduced one
41 of the most important dimensionless parameters in fluid mechanics, the ratio of momentum over
42 viscous forces which is called Reynolds number ever since. Based on this dimensionless parameter, it

43 was observed that irrotationality in the streamlines occurred for values much greater than 1 and led to
 44 somehow confine the occurrence of turbulence to Reynolds number values greater than
 45 approximately 1000 to 2000. Richardson (1922) introduced the idea of turbulence ‘energy cascade’ by
 46 stating that turbulent motion, powered by the kinetic energy, is first produced at the largest scales
 47 (through eddies of size comparable to the characteristic length scale of the natural process) and then to
 48 smaller and smaller ones, until is dissipated by the viscous strain action. Taylor (1935) was the first to
 49 use stochastic tools to study this phenomenon modelling turbulence by means of random variables
 50 rather than deterministic ones. Following this idea, Kolmogorov (1941a-c) managed to derive the
 51 famous ‘5/3’ law (K41 theory) using the Navier-Stokes equations. That law describes the energy
 52 dissipation rate from larger to smaller turbulence scales within the inertial wavenumber sub-range,
 53 with the power spectrum no longer dependent on the eddy size and fluid viscosity. Since then, many
 54 scientists (including Von Karman 1948, Heisenberg 1948, Kraichnan 1959, Batchelor 1959, Pope 2000),
 55 have significantly contributed to the current power-spectrum-based models of turbulence.

56 A general view of the stochastic approach of stationary and isotropic turbulence (in which the random
 57 variables describing turbulence have the same statistical properties in all directions) can be seen in
 58 many text books, e.g. Pope (2000). In this paper, we focus on the investigation of the second-order
 59 statistics (e.g. power spectrum) and the preservation of the marginal probability density function
 60 (pdf). We are mainly interested in the local and global stochastic properties of a process, by calculating
 61 its fractal dimension and by examining whether it exhibits HK behaviour, respectively. Furthermore,
 62 we investigate the stochastic properties of the most common three-dimensional power-spectrum-
 63 based models of stationary and isotropic turbulence in time domain and we detect some model
 64 weaknesses despite their widespread use. A simpler and more robust model, which incorporates both
 65 fractal and Hurst-Kolmogorov (HK) possible behaviours, is proposed using a second-order stochastic
 66 framework based on the concept of climacogram. This model is tested over high resolution nearly
 67 isotropic observational data of laboratory scale. Moreover, we show that the same model can be used
 68 for small-scale hydrometeorological processes generated by turbulence such as atmospheric wind
 69 speed and precipitation intensities. Expressions of other stochastic tools such as the autocovariance
 70 and power spectrum are also produced directly from the model and are in agreement with data.
 71 Finally, uncertainty, discretization and bias related errors are estimated for each stochastic tool,
 72 showing, in general, lower errors for the climacogram-based model and larger ones for power-
 73 spectrum based ones. It is noted that the HK process corresponds to Fractional Gaussian Noise (cf.
 74 Mandelbrot and Wallis 1968) and is named after Hurst (1951), who first detected the long-term
 75 behaviour in geophysical time-series and Kolmogorov (1940) who first introduced the mathematical
 76 form of the process (cf. Koutsoyiannis 2011a).

77 2. Definitions and notations

78 Stochastic modelling and probabilistic approaches have been proven useful in the investigation of
 79 processes that resist a deterministic description, such as turbulence (e.g. Kraichnan 1991 ch. 1, Frisch,
 80 2006 ch. 3, McDonoug, 2007 ch. 1, Koutsoyiannis 2014). Using stochastic mathematical processes one
 81 can represent, and thus interpret, a natural process based on its statistical properties whose values can
 82 be estimated through stochastic tools such as autocovariance-based ones defined in the equations
 83 below:

$$84 \quad c(\tau) := \text{Cov}[\underline{x}(t), \underline{x}(t + \tau)] \quad (1)$$

$$85 \quad v(\tau) := c(0) - c(\tau) \quad (2)$$

$$86 \quad s(w) := 4 \int_0^\infty c(\tau) \cos(2\pi w\tau) d\tau \quad (3)$$

87 where $\underline{x}(t)$ is the continuous time process (underscore denotes a random variable), $c(\tau)$ is the
 88 autocovariance function, $v(\tau)$ the variogram (else known as 2nd structural function), $s(w)$ the power
 89 spectrum and τ , w the continuous time lag and frequency, respectively (see in Appendix for details).

90 Other stochastic tools can be based on the climacogram (e.g. Koutsoyiannis 2013a), which is defined as
 91 the (plot of) variance of the averaged process $\frac{1}{m} \int_0^m \underline{x}(t) dt$ (assumed stationary) *vs* averaging time scale
 92 m and is denoted as $\gamma(m)$:

$$93 \quad \gamma(m) := \frac{\text{var}[\int_0^m \underline{x}(t) dt]}{m^2} \quad (4)$$

94 The climacogram is useful to measure the variance of a process among scales (the kinetic energy, in
 95 case the variable under consideration is the velocity), and has many advantages in stochastic model
 96 building, namely small statistical as well as uncertainty errors (Dimitriadis and Koutsoyiannis 2015). It
 97 is also directly linked to the autocovariance function by the following equations (Koutsoyiannis
 98 2013a):

$$99 \quad \gamma(m) = 2 \int_0^1 (1-x)c(xm) dx \quad (5)$$

$$100 \quad c(\tau) = \frac{\partial^2(\tau^2 \gamma(\tau))}{2\partial\tau^2} \quad (6)$$

101 A climacogram-based spectrum (CBS), else known as the ‘pseudospectrum’, for comparison with the
 102 classical power spectrum, can be also defined as (Koutsoyiannis 2013a):

$$103 \quad \psi(m) := \frac{2\gamma(1/w)}{w} \left(1 - \frac{\gamma(1/w)}{\gamma(0)}\right) \quad (7)$$

104 Furthermore, we introduce here, a climacogram-based variogram (CBV) for comparison with the
 105 classical variogram:

$$106 \quad \xi(m) := \gamma(0) - \gamma(m) \quad (8)$$

107 Note that both CBS and CBV include the process variance at scale 0, i.e. $\gamma(0)$ and thus, they are
 108 applied only after a stochastic model is set.

109 All the above stochastic tools definitions and expressions in discrete time as well as widely used
 110 estimators, estimations (based on the latter estimators) and expected values, can be found in
 111 Appendix.

112 **3. Most common stochastic models of stationary and isotropic** 113 **turbulence**

114 It is noted that the log-log derivative (LLD) is an essential concept in turbulence as it can identify
 115 possible scaling behaviour related to asymptotic coefficients (e.g. fractal dimension and Hurst
 116 coefficient). The LLD of any function $f(x)$ is defined as:

$$117 \quad f^\#(x) := \frac{d \ln(f(x))}{d \ln x} = \frac{x}{f(x)} \frac{df(x)}{dx} \quad (9)$$

118 and for the finite logarithmic derivative of $f(x)$, e.g. in case of discrete time process, we choose the
 119 backward log-log derivative, i.e.:

120 $f^\#(x_i) := \frac{\ln(f(x_i)/f(x_{i-1}))}{\ln(x_i/x_{i-1})}$ (10)

121 Based on Gneiting et al. (2012) analysis, the fractal dimension (F) can be defined as (cf. Beran et al.
122 2013 ch. 3.6):

123 $F := N + 1 - \frac{1}{2} \lim_{\tau \rightarrow 0} \xi^\#(\tau)$ (11)

124 where N the dimension of the field (e.g. $N=1$ for 1D velocity field).

125 Based on Beran et al. (2013 ch. 1.3) analysis, the Hurst coefficient (H) can be defined as:

126 $H := 1 + \frac{1}{2} \lim_{m \rightarrow \infty} \gamma^\#(m)$ (12)

127 3.1 Commonly used processes

128 Following the stochastic framework in Section 2 (and in Appendix), we derive in Table 1, the 1D and
129 3D isotropic power spectra as well as their LLD's, for a Markovian process, a special case of a
130 powered-exponential process (e.g. Yaglom 1987 ch. 10, Gneiting et al. 2012) and a generalized HK
131 (gHK) process (cf. Dimitriadis and Koutsoyiannis 2015), which the latter behaves as Markovian-like
132 for small scales and HK-like for large ones. These positively-correlated mathematical processes
133 enclose possible asymptotic behaviours in large and small scales. In particular, a positively-correlated
134 natural process may approach zero or infinite scale, by a powered-exponential (e.g. Markovian
135 process) or a power-type (e.g. HK process) rise or decay, respectively. The 1D power spectrum and the
136 3D one, denoted as $s_{3D}(\mathbf{w})$, are related by (Batchelor 1959 p. 50, Pope 2000 pp. 226-227, Kang et al.
137 2003):

138 $s(w) = \int_1^\infty \frac{x^2-1}{x^3} s_{3D}(\|\mathbf{w}\|x) dx$ (13)

139 $s_{3D}(w) = \frac{w^3}{2} \frac{\partial \left(\frac{1}{w} \frac{\partial(s(w))}{\partial w} \right)}{\partial w}$ (14)

140 where \mathbf{w} is the isotropic 3D frequency vector, with $\|\mathbf{w}\| = w \geq 0$.

141 As mentioned above, the most common used model for stationary and isotropic turbulence consists of
142 the work of many scientists. Combining them into one equation, the power spectrum of isotropic and
143 stationary turbulence can be written as (Pope 2000 pp. 232-233, Cerutti and Meneveau 2000, Kang et
144 al. 2003):

145 $s_{3D}(w) = f_E(w, c_E, p) f_I(w, c_I) f_D(w, c_D)$ (15)

146 where c_E , c_I , c_D and p are model parameters (see Pope 2000, pp. 233 for description) and from the
147 work of Von Karman (1948), for the from the work of Von Karman (1948), for the energy containing
148 eddies (large scales):

149 $f_E(w, c_E, p) = \left(\frac{w}{\sqrt{w^2 + c_E}} \right)^{\frac{5}{3} + p}$ (16)

150 combined with the work of Kolmogorov (1941a-c) for the inertial range (intermediate scales):

151 $f_I(w, c_I) = c_I w^{-\frac{5}{3}}$ (17)

152 and from the work of Kraichnan (1959) for the dissipation range (small scales):

153 $f_D(w, c_D) = e^{-w c_D}$ (18)

154 **Table 1:** 1D and 3D power spectrum for Markovian, powered-exponential and gHK processes as well
 155 as their LLD's (estimated from equation 9), where λ is the parameter related to the true variance of the
 156 process, q the scale parameter and b is related to the power-type behaviour of the process.

Markovian		Powered-exponential special case		gHK	
$c(\tau) = \lambda e^{- \tau /q}$	(19)	$c(\tau) = \lambda e^{-(\tau/q)^2}$	(20)	$c(\tau) = \lambda \frac{(1-b)(2-b)}{(1+ \tau /q)^b}$ with $b \in (0,2)$	(21)
$s(w) = \frac{4\lambda q}{1 + 4\pi^2 q^2 w^2}$ with $\lim_{w \rightarrow 0} s^\# = 0$ and $\lim_{w \rightarrow \infty} s^\# = -2$	(22)	$s(w) = \frac{\lambda q \sqrt{\pi}}{2} e^{-(qw\pi)^2}$ with $s^\#(w) = -2(qw\pi)^2$, $\lim_{w \rightarrow 0} s^\# = 0$ and $\lim_{w \rightarrow \infty} s^\# = -\infty$	(23)	$\lim_{w \rightarrow 0} s \sim w^{b-1}$ with $\lim_{w \rightarrow 0} s^\# = b - 1$	(24)
				$\lim_{w \rightarrow \infty} s \sim w^{-2}$ with $\lim_{w \rightarrow \infty} s^\# = -2$	(25)
$s_{3D}(w) = \frac{4\lambda q (2\pi q w)^4}{(1 + 4\pi^2 q^2 w^2)^3}$ with $\lim_{w \rightarrow 0} s_{3D}^\# = 4$ and $\lim_{w \rightarrow \infty} s_{3D}^\# = -2$	(26)	$s_{3D}(w) \sim q^5 w^4 e^{-(qw\pi)^2}$ with $s^\#(w) = 4 - 2(qw\pi)^2$ $\lim_{w \rightarrow 0} s_{3D}^\# = 4$ and $\lim_{w \rightarrow \infty} s_{3D}^\# = -\infty$	(27)	$\lim_{w \rightarrow 0} s_{3D} \sim w^{b-1}$ with $\lim_{w \rightarrow 0} s_{3D}^\# = b - 1$	(28)
				$\lim_{w \rightarrow \infty} s_{3D} \sim w^{-2}$ with $\lim_{w \rightarrow \infty} s_{3D}^\# = -2$	(29)

157

158 3.2 Stochastic properties of large-scale range

159 For the 3D and 1D (derived from the 3D one) power spectra at the energy containing range, we have
 160 that:

161 $\lim_{w \rightarrow 0} s_{3D} = \lim_{w \rightarrow 0} s \sim w^p$ (30)

162 where Von Karman (1948) suggests $p = 4$ (or else known as 'Batchelor turbulence', cf. Davidson 2000),
 163 while other works result in different values, e.g. Saffman (1967) suggests $p = 2$.

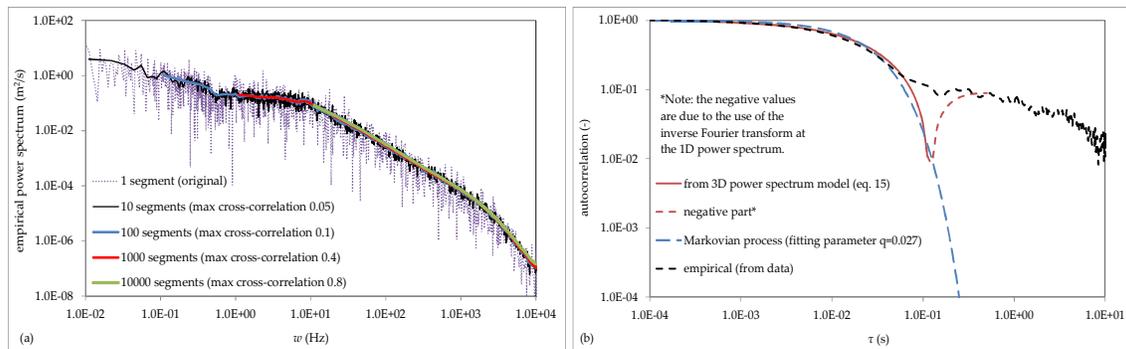
164 There are many arguments about the proper value of the p parameter and its relation to the
 165 Loitsyansky integral which controls the rate of decay of kinetic energy (cf. Davidson 2000). The main
 166 debate is whether points at a large distance in stationary, isotropic and homogeneous turbulent flow
 167 are statistically independent or show a correlation that decays either exponentially (e.g. Von Karman
 168 model for wind gust, cf. Wright and Cooper 2007 ch. 16.7.1; Faisst and Eckhardt 2004, Avila et al. 2010
 169 and Kuik et al. 2010, models for pipe flow) or with a power-type law (see below for several examples).

170 Towards the stochastic properties of the aforementioned equation, we can see from Table 1 that the
 171 case $p = 2$ does not correspond neither to exponential (Markovian or powered-exponential) nor to
 172 power-type (i.e. HK) decay of autocovariance. Hence, this model cannot be applied to asymptotic zero

173 frequencies (or infinite scales). Interestingly, the case $p = 4$ can be interpreted by a Markovian
 174 (equation 26) or a special case of the powered-exponential (equation 27) decay of autocovariance.
 175 However, this case also excludes the HK behaviour, i.e. autocovariance long-range dependence (e.g.
 176 equation 21), where p now equals $b - 1$ and is bounded to $[-1, 1]$.

177 Although the aforementioned models do not include a possible power-law decay of autocovariance
 178 (i.e. HK behaviour), several works show strong indication that turbulence natural processes can
 179 exhibit HK behaviour rather than Markovian. Such works are reported by e.g., Nordin et al. (1972) for
 180 laboratory turbulent flume and turbulent river velocities, Helland and Van Atta (1978) for grid
 181 turbulence velocities, Goldstein and Roberts (1995) for magneto-hydrodynamic turbulent solar wind,
 182 Chamorro and Potre-Agel (2009) for wind turbulent wakes and grid-turbulence, Dimitriadis and
 183 Papanicolaou (2012) and Charakopoulos et al. (2014a,b) for turbulent buoyant jets, Koutsoyiannis
 184 (2013b) for grid turbulence. Koutsoyiannis (2011b) has also shown that entropy maximization results
 185 in HK dynamics at asymptotic times (zero or infinity) under the constraints of mean, variance and
 186 autocovariance of lag one preservation.

187 We believe that the reason a possible HK behaviour is not detected in geophysical processes (which
 188 are often characterized by lack of measurements), is that mathematical smoothing techniques are
 189 applied (e.g. windowing or else Welch approaches, regression analysis, wavelet techniques, see other
 190 examples in Stoica and Moses 2004 ch. 2.6). Particularly, application of windowing techniques to any
 191 stochastic tool can be misleading since they eliminate a portion (depending on the type and length of
 192 the window applied) of the time-series' variance (which often is incorrectly attributed to 'noise', cf.
 193 Koutsoyiannis, 2010). This elimination can lead to process' misrepresentation in case of significant
 194 effects of discretization, small and/or finite record length and bias (examples of applications to the
 195 power spectrum can be seen in e.g. Lombardo et al. 2013). An example of smoothing out the HK
 196 behaviour by applying the Welch approach with a Bartlett window and no segment-overlapping to an
 197 observed time-series, is shown in Fig. 1(a). Even though the smoothing technique decreases the power
 198 spectrum variance, it also causes low frequency loss of information (e.g. see other examples in
 199 Dimitriadis et al., 2012). This loss of information may cause a process misinterpretation, as illustrated
 200 in Fig. 1(b), where the 1D autocorrelation function (derived from the 3D power spectrum model in
 201 equation 15) exhibits a Markovian-like decay, while the empirical one (derived from the windowed
 202 empirical power spectrum partitioned into 10^3 segments) exhibits an HK behaviour. Also, this
 203 smoothing technique should be used in caution in strong-correlated processes, as increasing the
 204 number of partitioned segments will also cause an increase in their cross-correlation (Fig. 1a). Finally,
 205 processes with HK behaviour have usually large bias and in case this is not included in the model, the
 206 empirical autocovariance's rapid decay in large scales (or equivalently lags) may be erroneously
 207 interpreted as short-range dependence (Fig. 1b).



208 **Fig. 1:** (a) Example of loss of low frequency information caused by the application of the windowing
 209 technique, in a time-series provided by the Johns Hopkins University (see also in Section 4 for more
 210 details on the dataset) as well as the maximum cross correlations between the partitioned segments;
 211

212 (b) 1D autocorrelation function derived from the 3D power spectrum model in equation 15 (with
 213 parameters based on the fitting of the windowed 1D power spectrum with 1000 segments in Fig. 1(a):
 214 $c_E = 2.5 \text{ m}^{-2}$, $p = 4$, $c_I = 13.0 \text{ m}^3/\text{s}^2$, $c_D = 2 \times 10^{-4} \text{ m}$); a Markovian autocorrelation function, i.e.
 215 $e^{-(\tau/q)}$, for reasons of comparison; and the corresponding (to the windowed 1D power spectrum with
 216 1000 segments in Fig. 1a) empirical autocorrelation function.

217 To incorporate possible HK behaviour in the model, we may assume an autocovariance power-type
 218 decay at large scales, where the 3D and 1D power spectra at asymptotically zero frequency are of the
 219 form w^{b-1} (Table 1), with b bounded to $(0,2)$, for positively correlated processes $(0.5 < H < 1)$,
 220 negatively-correlated processes $(0 < H < 0.5)$ and for a process with a random decay in large scales
 221 $(H = 0.5)$, with H the Hurst coefficient $(H = 1 - b/2)$, from equation 12).

222 3.3 Stochastic properties of small-scale range

223 Similarly, for the 3D and 1D power spectra at the dissipation range, we have that:

$$224 \lim_{w \rightarrow \infty} S_{3D}(w) = \lim_{w \rightarrow \infty} s(w) \sim e^{-w} \quad (31)$$

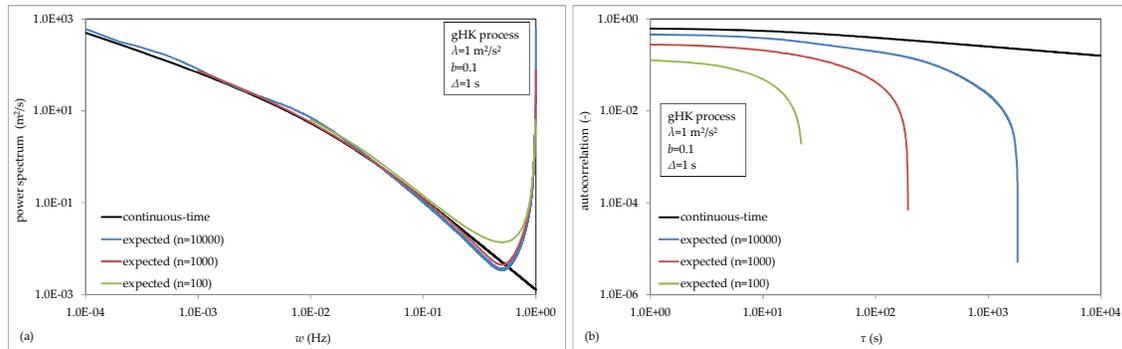
225 This results in autocovariance function of the form:

$$226 c(\tau) \sim \frac{1}{\tau^2 + 1} \quad (32)$$

227 which corresponds to Wackernagel (1995) process (he also refers to it as autocovariance-based
 228 Cauchy-class process resembling the Cauchy probability function). A generalized expression of this
 229 process can be found in Gneiting (2000), which we will refer to it as the Gneiting process (its analytical
 230 expressions are shown in Section 4.2). For small lags this process behaves like (e.g. Gneiting and
 231 Schlather 2004):

$$232 \lim_{\tau \rightarrow 0} c(\tau) \sim 1 - \tau^2 \sim e^{-\tau^2} \quad (33)$$

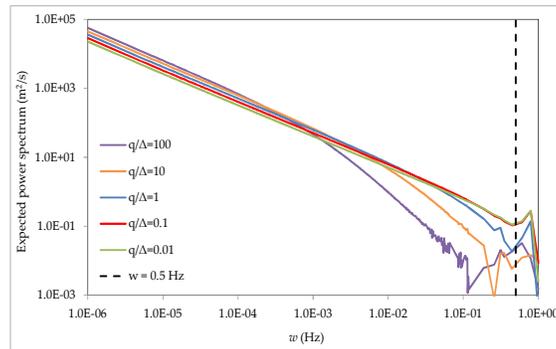
233 which corresponds to the special case of a powered-exponential process in Table 1. Note, that this
 234 process corresponds to $H = 0$ (based on the definition in equation 12), if applied to large scales.



235 **Fig. 2:** (a) Power spectra and (b) corresponding autocovariances, in continuous time as well as their
 236 expected values, with varying number of records (denoted as n) of a gHK process. The expected
 237 autocovariance and power spectrum are estimated from equation (A17) and (A25), respectively (see
 238 Appendix).
 239

240 Other models for the dissipation range are of the form of a powered-exponential power spectrum
 241 process (e.g. Cerutti and Meneveau 2000) which may result from a powered-exponential
 242 autocovariance function (Table 1). However, there is evidence that these models cannot interpret the
 243 frequently observed spike in the high frequency power spectrum (e.g. Cerutti and Meneveau 2000,

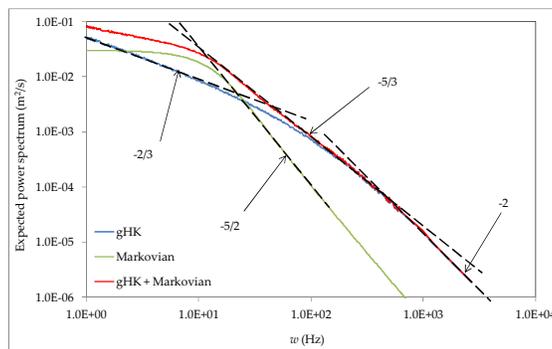
244 Kang et al. 2003). This is usually ignored and attributed to instrumental noise. Here, we show that this
 245 spike may appear in HK processes and is due to discretization and bias errors, in case the shape
 246 parameter q/Δ takes large values (Fig. 3).



247
 248 **Fig. 3:** Expected power spectra (estimated from equation A25) of a gHK process, with varying q/Δ
 249 (where Δ the sampling time interval, see in Appendix for its relation to the expected value of a
 250 stochastic tool).

251 3.4 Stochastic properties of intermediate-scale range

252 From Table 1, one may observe that the power spectrum asymptotic LLD's from different processes,
 253 are often coincident with each other. For example, for both a Markovian and a gHK process with $b=1$,
 254 the power spectrum LLD is 0 for the low frequency tail and -2 for the high frequency one. This may be
 255 confusing and result in misinterpretation of the natural process. A solution to this may be to
 256 incorporate additional stochastic tools in the analysis as shown in Section 4. For the aforementioned
 257 example, if the autocovariance function asymptotic properties (local and global ones) are analyzed,
 258 one can decide upon a powered-exponential lag decay (e.g. a Markovian process) and a power-type
 259 one (e.g. a gHK process). At the same basis, when a power-type behaviour appears in the intermediate
 260 frequencies of a power spectrum (e.g. in case of a $-5/3$ LLD), it may be misleading to interpret it as a
 261 power-law function (and thus, a power-type autocovariance decay, as shown in Table 1), because this
 262 can result from different kind of processes which they do not have power-type expressions for the
 263 intermediate scale-range. An illustrative example is shown in Fig. 4, where the $-5/3$ LLD in the
 264 intermediate frequencies of the power spectrum results from a simple combination of a Markovian
 265 and a gHK process, both of which have a purely stochastic interpretation and they do not include
 266 power-type laws in the intermediate frequency-range.



267
 268 **Fig. 4:** Expected power spectrum (estimated from equation A25) resulted from a combination of a
 269 Markovian and a gHK process (with parameters same as in the application of section 4.1 and $N=10^4$).

270 Note also, that the Kolmogorov (1941a-c) power-type power spectrum refers only to intermediate
 271 frequencies and should not be applied arbitrarily for low frequencies too, as the corresponding

272 autocovariance asymptotic large-scale behaviour, i.e. $c(\tau) \sim \tau^{\frac{5}{3}-1}$, gives an invalid (based on equation
 273 12) $H = 4/3 > 1$.

274 4. Proposed model and applications

275 In the previous section, we present several limitations concerning the stochastic properties of
 276 proposed turbulent models from literature. Specifically, we see that they only include exponential
 277 decay in the energy containing area and thus, completely excluding possible HK behaviour. They also,
 278 describe the dissipation area decay with only a specific case of a powered-exponential process and
 279 thus, leaving out all other possible types of decay. Moreover, they interpret a possible power-type-like
 280 intermediate area (of the power spectrum) with power-type behaviour (and particularly, only that of
 281 the K41 theory) which can also result from intermediate non power-type processes (as shown in Fig.
 282 4). Furthermore, these models are based only on the power spectrum stochastic tool (causing possible
 283 misinterpretation in other tools, e.g. climacogram, autocovariance) and on multiple processes
 284 multiplication (which may cause numerical difficulties in stochastic generation). Since turbulence
 285 generates and drives most of geophysical processes, we expect geophysical processes to exhibit similar
 286 types of decay in small and large scales. Hence, a more robust, flexible and parsimonious model is
 287 required that can incorporate all the aforementioned microscale and macroscale behaviours linking
 288 turbulence to hydrology. Here, we choose the ergodic stochastic model in Table 2, which consists of
 289 two independent processes, that of a powered exponential (controlling the small scales and fractal
 290 behaviour, cf. Gneiting et al. 2012) and a gHK (controlling the large scales and HK behaviour, cf.
 291 Dimitriadis and Koutsoyiannis, 2015), which are combined in a way to exhibit the desired expected
 292 LLD in the intermediate scales. This model can describe all linear combinations of powered-
 293 exponential and HK processes, including the often observed intermediate quick drop of all the
 294 stochastic tools (see Section 4.1, 4.2 and 4.3, for an example in grid turbulence, wind and precipitation
 295 process). This particular drop may be due to the interference of boundaries and/or the existence of
 296 multiple periodic functions, as for example in case of combinations of HK with cyclostationary
 297 processes (cf. Markonis and Koutsoyiannis 2013). Furthermore, although the proposed model results
 298 in a complicated power spectrum expression (equation 37), it provides simpler expressions for the
 299 other tools if compared to the most common model described in Section 3 (which has no analytical
 300 expressions for all tools except for the power spectrum). Finally, the proposed model is also justified
 301 by the maximization of entropy production in logarithmic time (abbreviated EPLT), a term introduced
 302 and defined by Koutsoyiannis (2011c) as the LLD of entropy. Particularly, Koutsoyiannis (2015)
 303 showed that the powered-exponential process has the largest EPLT for the microscale range (time-
 304 scale tending to zero) and the HK process has the largest EPLT for the macroscale range (time-scale
 305 tending to infinity). Hence, the maximization of EPLT can result from a combination of both
 306 processes.

307 **Table 2:** Autocovariance, variogram, climacogram, CBV, CBS and power spectrum mathematical
 308 expressions of the stochastic model, consisted of two independent processes in continuous time, that
 309 of a powered exponential and a gHK.

Type	Stochastic model
Autocovariance*	$c(\tau) = \lambda_1 e^{-(\tau /q_1)^a} + \lambda_2 (1 + \tau /q_2)^{-b}$ (34)

$$\text{Climacogram} \quad \gamma(m) = \frac{2\lambda_1 \left(\frac{m}{q_1} \Gamma_1(1/a, (m/q_1)^\alpha) - \Gamma_1(2/a, (m/q_1)^\alpha) \right)}{(m/q_1)^2 + \frac{2\lambda_2((m/q_2 + 1)^{2-b} - (2-b)m/q_2 - 1)}{(1-b)(2-b)(m/q_2)^2}} \quad (35)$$

$$\text{Variogram} \quad v(\tau) = \lambda_1 + \lambda_2 - c(\tau) \quad (36)$$

$$\text{Power spectrum}^{**} \quad s(w) = \text{ICF}[\lambda_1 e^{-(|w|/q_1)^a}] + \frac{4\lambda_2 q_2^b \Gamma(1-b) \text{Sin}\left(\frac{\pi b}{2} + 2q_2 \pi |w|\right)}{(2\pi |w|)^{1-b}} - \frac{4\lambda_2 q_2 {}_1F_2\left[1; 1 - \frac{b}{2}, \frac{3}{2} - \frac{b}{2}; -\pi^2 q_2^2 w^2\right]}{1-b} \quad (37)$$

$$\text{CBV} \quad \xi(m) = \lambda_1 + \lambda_2 - \gamma(m) \quad (38)$$

$$\text{CBS} \quad \psi(w) = \frac{2\gamma(w)}{w} \left(1 - \frac{\gamma(w)}{\lambda_1 + \lambda_2} \right) \quad (39)$$

310 * $\lambda_2 = \lambda(1-b)(2-b)$, with λ a parameter related strictly to the process' variance.

311 ** Since the inverse cosine Fourier (ICF) transform of the powered-exponential function and the hyper-
 312 geometric function ${}_1F_2$ have not an analytical form, this cannot be written in a closed expression and
 313 numerical algorithms must be used.

314 4.1 Application to small-scale grid turbulence

315 In this section, we show the stochastic analysis of a grid-turbulence process based on a large open
 316 access dataset (<http://www.me.jhu.edu/meneveau/datasets/datamap.html>), provided by the Johns
 317 Hopkins University. Microscale turbulence description has many applications in hydrometeorological
 318 processes which often lack small scale measurements (cf. Koutsoyiannis 2011c), thus introducing
 319 limitations in the fitted models (e.g. the fractal dimension of the process cannot be estimated based on
 320 the definition of equation 11). An illustrative example of an application to atmospheric wind speed is
 321 shown in Section 4.2.

322 Here, we only consider the longitudinal wind velocity dataset along the flow direction since the other
 323 two components are limited by the experiment's construction boundaries. This dataset consists of 40
 324 time-series (Fig. 5a), measured by X-wire probes placed downstream of the grid (Kang et al. 2003). The
 325 first 16 time-series correspond to velocities measured at transverse points abstaining $r = 20M$ from the
 326 source, where $M = 0.152$ m is the size of the grid. The next 4 time-series correspond to distance $r =$
 327 $30M$, the next 4 to $40M$ and the last 16M to $48M$. For details regarding the experimental setup and
 328 datasets see Kang et al. (2003). All time-series are considered to be stationary with a nearly-Gaussian
 329 probability density function (see in Fig. 5c), are nearly isotropic with isotropy ratio 1.5 (Kang et al.
 330 2003) and very long (each contains $n = 36 \times 10^6$ data points), covering all three aforementioned scale
 331 ranges of equation (15). Moreover, the sampling time interval, denoted as D , is considered small (2.5
 332 μ s), therefore equality $D = \Delta$, where $\Delta (\leq D)$ the instrument response time, can be assumed valid. In
 333 Appendix, we noted that if D is small the differences between stochastic processes in discretized time
 334 with $\Delta > 0$ and $\Delta \approx 0$ are also expected to be small. Finally, following the same analysis of Dimitriadis
 335 and Koutsoyiannis (2015), the expected value of each examined stochastic tool can be roughly
 336 estimated as the average value of all 40 time-series (Fig. 6a-g), after homogenization is applied (the
 337 marginal variance of the process is estimated approximately $2.272 \text{ m}^2/\text{s}^2$). Additionally, we choose the
 338 38th time-series for the empirical one, after observing that is the closest one to each stochastic tool's
 339 averaged value (Fig. 6h). Since we expect this to be near to the process expected values, it can help us
 340 test the validity of the stochastic model. Modelling phenomena such as intermittency (which is related

341 to high-order derivatives, c.f. Kang et al. 2003, Batchelor and Townsend 1949) as well as preservation
 342 of high order moments (which are often characterized by high uncertainty, cf. Lombardo et al. 2014)
 343 deviate from the purpose of this paper. In this paper, we are mainly interested in the local and global
 344 2nd order stochastic properties of the process, by calculating the process fractal dimension and by
 345 examining whether the process exhibits HK behaviour, respectively.

346 As we have already mentioned, the velocity field is not homogeneous and the root-mean-square (rms)
 347 velocity components (i.e. standard deviations of velocity) are decreasing with the distance from the
 348 grid (Fig. 5b). To make data homogeneous, we normalize each time-series by subtracting the mean
 349 $\mu_t(r)$ and dividing by the standard deviation $\sigma_t(r)$, both estimated from the equations of the fitted
 350 curves in Fig. 5(b):

$$351 \quad \sigma_t(r) = 4.16(r + 0.3)^{-0.657} \quad (40)$$

$$352 \quad \sigma_t(r)/\mu_t(r) = 0.859r + 3.738 \quad (41)$$

353 where r is the distance from the grid. Note that coefficient 0.3 in equation (40) has been added for
 354 consistency reasons, so that the variance is finite at distances near the grid.

355 We also observe that the pdf of the time-series are not exactly Gaussian, since for example the
 356 empirical skewness is approximately equal to 0.2 (Fig. 5c and 5d). Here, we propose a normalization
 357 scheme by separating the empirical pdf to multiple segments and then approximating them with
 358 multiple Gaussian distributions:

$$359 \quad f_t(u) = \begin{cases} N(\mu_1, \sigma_1), & -\infty < u_1 \leq h_1 \\ N(\mu_2, \sigma_2), & h_1 < u_2 \leq h_2 \\ \dots & \dots \\ N(\mu_o, \sigma_o), & h_{o-1} < u_o < \infty \end{cases} \quad (42)$$

360 where $f_t(u)$ is the model pdf of the velocity u , $N(\mu_l, \sigma_l)$ is a Gaussian pdf for the u_l branch of the
 361 empirical pdf (consisted of all quantiles $h_{l-1} < u_l \leq h_l$), with l varying from 1 to o (with $h_0 \rightarrow -\infty$ and
 362 $h_o \rightarrow \infty$) and with o representing the number of branches we separate the empirical pdf.

363 The μ_l and σ_l parameters can be calculated by simply fitting $N(\mu_l, \sigma_l)$ to the empirical pdf of the
 364 quantiles within the l segment (subject to the constraints that the cdf and pdf values between the
 365 multiple Gaussian functions are equal). Specifically, if the l segment consists of only two quantiles, u_1
 366 and u_2 , and with F_1 and F_2 , the empirical cumulative distribution function (cdf) at these points, then
 367 the above parameters are obviously equal to:

$$368 \quad \mu_l = u_1 - \sigma_l \sqrt{2} \operatorname{erf}^{-1}(2F_1 - 1) \quad (43)$$

$$369 \quad \sigma_l = \frac{u_2 - u_1}{\sqrt{2}(\operatorname{erf}^{-1}(2F_2 - 1) - \operatorname{erf}^{-1}(2F_1 - 1))} \quad (44)$$

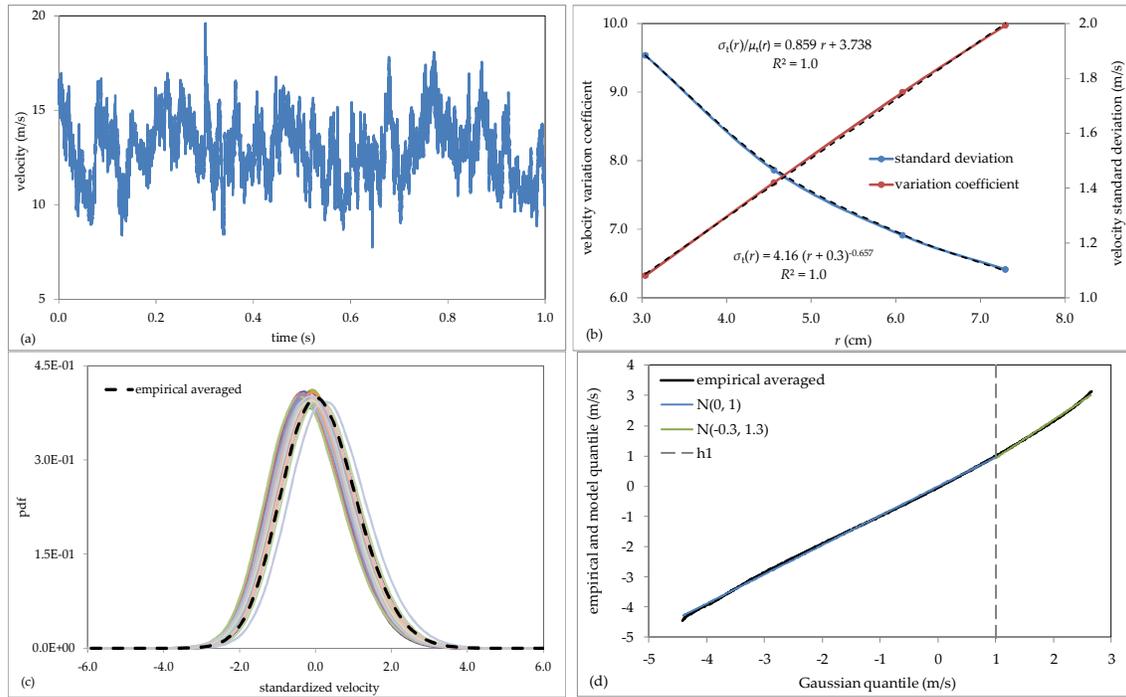
370 with erf^{-1} the inverse of the error function.

371 Then, we can easily transform $u \sim f_t$ to $u_n \sim N(0,1)$, by simply subtracting from each set of quantiles
 372 ($h_{l-1} < u_l \leq h_l$) the mean μ_l and then dividing with the standard deviation σ_l . Furthermore, the
 373 reverse transformation scheme from a variable $u_n \sim N(0,1)$ to $u_r \sim f_t$, can be easily done by multiplying
 374 each set of quantiles ($h'_{l-1} < u_{n,l} \leq h'_l$) from u_n , with σ_l and then by adding μ_l (where $h'_{l-1} = \frac{h_{l-1} - \mu_l}{\sigma_l}$,
 375 $h'_l = \frac{h_l - \mu_l}{\sigma_l}$ and $u_{n,l} = \frac{u_l - \mu_l}{\sigma_l}$). This scheme can be easily applied to any type of empirical pdf, however in
 376 cases where the empirical pdf highly deviates from a Normal pdf, a large number of segments may be
 377 acquired and the process' pdf be poorly interpreted.

378 Here, we observe that the left and right branch of the averaged empirical pdf can be very well
 379 approximate by two Gaussian distributions. Thus, we approximate the pdf of the process with 2
 380 segments ($\nu = 2$), with parameters shown in Fig. 5(b), with Pearson correlation coefficient $R^2 = 0.995$,
 381 between the empirical and the modelled pdf of equation (45):

$$382 \quad f_t(u) = \begin{cases} N(0,1), & -\infty < u \leq 1 \\ N(-0.3,1.3), & 1 < u < \infty \end{cases} \quad (45)$$

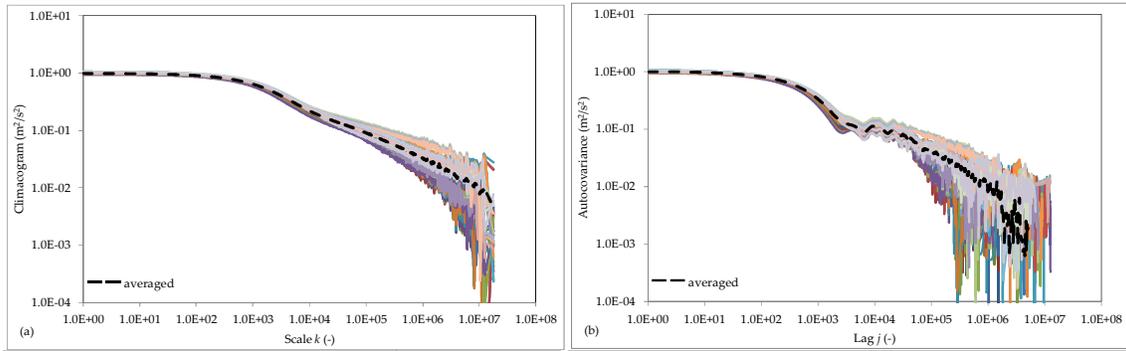
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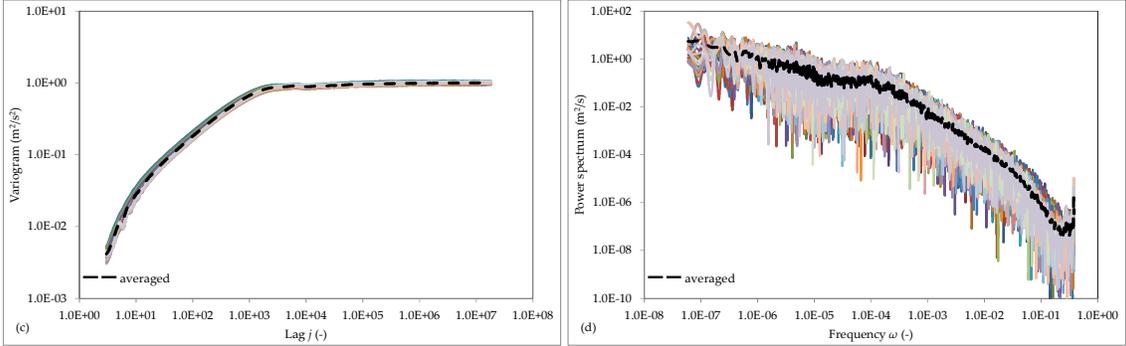
384 **Fig. 5:** Data preliminary analysis: (a) 1 s time window of one of the raw time-series; (b) averaged
 385 velocity mean $\mu_t(r)$ divided by the averaged velocity standard deviation $\sigma_t(r)$ (variation coefficient)
 386 and averaged velocity standard deviation $\sigma_t(r)$ as a function of r , along the longitudinal axis, as well
 387 as their fitted curves (black dashed lines); (c) empirical pdf's of the standardized time-series (multi-
 388 coloured lines) by subtracting $\mu_t(r)$ and dividing with $\sigma_t(r)$ each time-series and the empirical
 389 averaged pdf; (d) qq-plot of averaged empirical pdf vs standard Gaussian pdf, i.e. $N(0,1)$, along with
 390 modelled pdf from equation 45 (all parameters in m/s).
 391

392 In Fig. 6, we show the climacograms, autocovariances, variograms, power spectra, CBV's and CBS's
 393 from all 40 standardized time-series, their averaged values and the corresponding values of the 38th
 394 time-series. Assuming that these averaged values are near the process' expected ones, we can fit a
 395 stochastic model based on all the stochastic tools examined, and particularly the ones with the
 396 smallest statistical error for each scale, lag and frequency. We observe (Fig. 6g-h) that the large scale
 397 autocovariance and climacogram expected LLD's are both larger than -1 and that the power spectrum
 398 and CBS low frequency expected LLD's are larger than 0. Hence, it is most probable that the process
 399 exhibits HK behaviour.

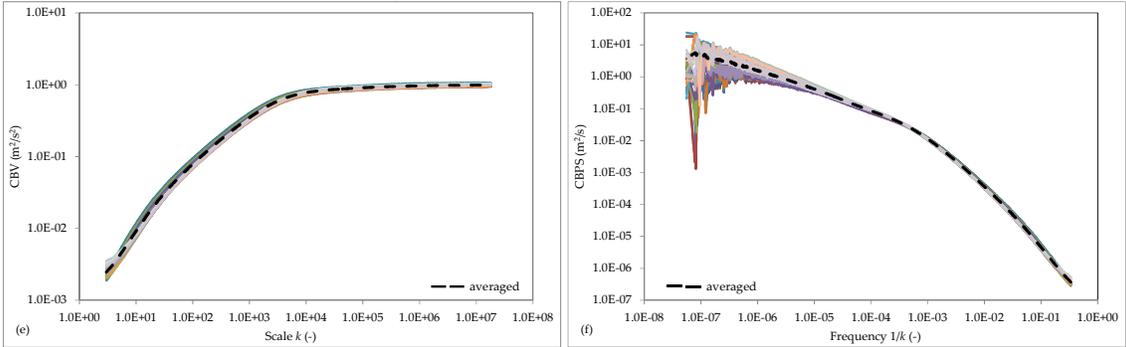
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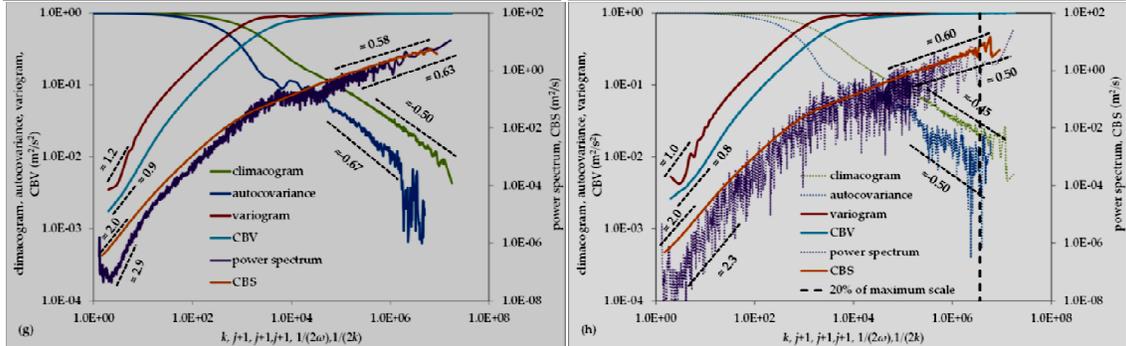


Fig. 6: Data analysis: (a) climacograms, (b) autocovariances, (c) variograms, (d) power spectra, (e) CBV and (f) CBS (with $\gamma(0)$ taken from the model in Table 2), of all the 40 time-series (multi-coloured lines) as well as their averaged values (black dashed lines); (g) all averaged values along with their averaged LLD's at large scales, lags and inverse frequencies and (h) those of the 38th time-series. Note that we use scales, lags and inverse frequencies up to the 20% of the maximum scale for our calculations, following the rule of thumb proposed in Koutsoyiannis (2003), Dimitriadis and Koutsoyiannis (2015).

Primarily, we try to best fit the climacogram-based stochastic tools (for reasons that will be explained later) and secondarily, the variogram for the intermediate lags (see in Fig. 7). To estimate the process

412 parameters in Table 2, a dimensionless fitting error is considered (as in Dimitriadis and Koutsoyiannis
413 2015):

$$414 \quad \text{FE}_\theta = \sum_z \left(\frac{\text{E}[\hat{\theta}(z)] - \hat{\theta}_d^{(A)}(z)}{\text{E}[\hat{\theta}(z)]} \right)^2 \quad (46)$$

415 where $\hat{\theta}_d^{(A)}$ is the empirical stochastic tool estimated from the data, $\text{E}[\hat{\theta}]$ the expected one estimated
416 from the model and z the corresponding to the stochastic tool scale, lag or frequency.

417 The optimization analysis results in scale parameters $\lambda_1 = 0.422 \text{ m}^2/\text{s}^2$ and $\lambda_2 = 0.592 \text{ m}^2/\text{s}^2$, shape
418 parameters $q_1 = 19.6 \text{ ms}$ and $q_2 = 1.45 \text{ ms}$, fractal parameter $a = 1.4$ and HK parameter $b = 0.32$,
419 with correlation coefficient R^2 approximating 1.0 for the climacogram and CBV, 0.99 for the CBS and
420 variogram, 0.95 for the autocovariance and 0.8 for the power spectrum.

421 Applying the L'Hôpital's rule and through mathematical calculations, we find that the fractal
422 dimension of the process in Table 2 is affected only by the exponent α of the powered-exponential
423 process and the Hurst coefficient only by the exponent b of the gHK one. Thus, process' fractal
424 dimension and Hurst coefficient are estimated (based on the definition in equation (11) and (12) and
425 Gneiting and Schlather 2004, analysis) as:

$$426 \quad F = 2 - \frac{\alpha}{2} = 1.3 \quad (47)$$

$$427 \quad H = 1 - \frac{b}{2} = 0.84 \quad (48)$$

428 Finally, to test the validity of our initial assumption, that for the specific model in Table 2 and the
429 estimated parameters the classical estimators of the climacogram-based stochastic tools have the
430 smallest error ε if compared to the autocovariance, variogram and power spectrum ones, we proceed
431 as follows. We calculate the statistical error for each stochastic tool via Monte Carlo analysis (since we
432 lack analytical expressions for the variance of the expected values):

$$433 \quad \varepsilon_\theta = \frac{\text{E}[(\hat{\theta} - \theta)^2]}{\theta^2} = \varepsilon_{\theta,v} + \varepsilon_{\theta,b} \quad (49)$$

434 where we have decomposed the dimensionless mean square error into a variance and a bias term (see
435 in Dimitriadis and Koutsoyiannis 2015),

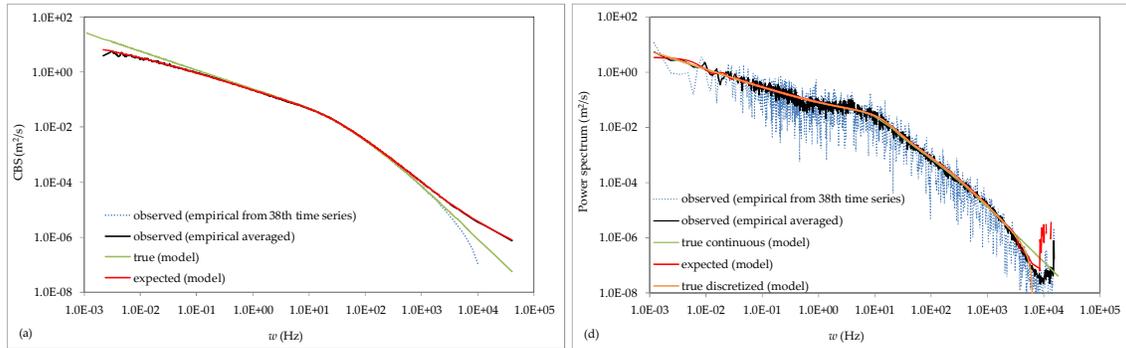
$$436 \quad \varepsilon_{\theta,v} = \text{var}[\hat{\theta}]/\theta^2 \quad (50)$$

$$437 \quad \varepsilon_{\theta,b} = (\theta - \text{E}[\hat{\theta}])^2/\theta^2 \quad (51)$$

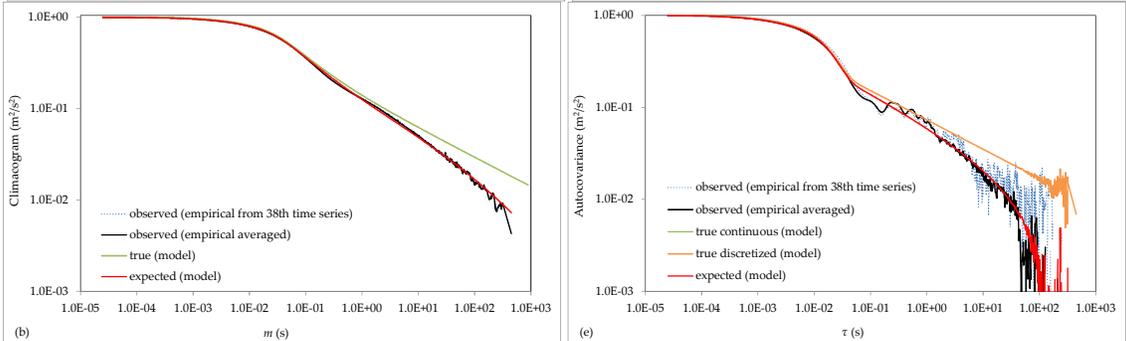
438 where θ is the examined stochastic tool, $\varepsilon_{\theta,b}$ can be easily estimated from equations in Tables A1-A6
439 and $\varepsilon_{\theta,v}$ is calculated from the Monte Carlo analysis since we lack analytical expressions.

440 Thus, we produce 40 time-series with $n = 36 \times 10^6$ using the SMA algorithm (Koutsoyiannis 2000 and
441 2015), which can replicate any stochastic process. Then, we compare the errors ε for each stochastic
442 tool for 81 points logarithmically distributed from 1 to n (Fig. 8). Note that in Fig. 8, we try to show all
443 estimates within a single plot for comparison. The inverse frequency in the horizontal axis is set to
444 $1/(2\omega)$, in order to vary between 1 and $n/2$, and the lag to $j+1$, so as the estimation of variance at $j = 0$ is
445 also shown in the log-log plot. From the results of this analysis, it can be observed that the initial
446 choice of the climacogram-based stochastic tools (and the variogram's for a small window of
447 intermediate LLD's) to interpret the empirical process, is proven valid for the current model structure,
448 model parameters and examined range of scales, with the power spectrum exhibiting the largest
449 errors.

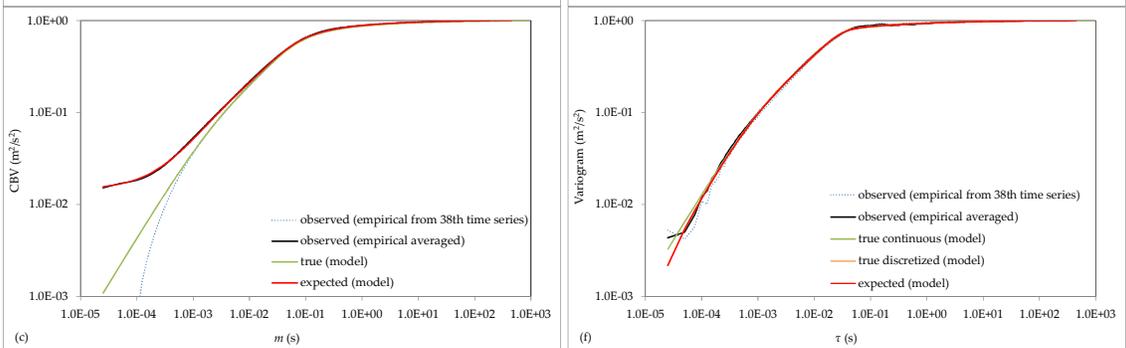
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Fig. 7: Stochastic modelling: true in continuous time (estimated from the model), true in discrete time (estimated from the model), expected (estimated from the model), empirical averaged (estimated from all 40 time-series) and observed (estimated from the 38th time-series), for the (a) CBS, (b) climacogram, (c) CBV, (d) power spectrum, (e) autocovariance and (f) variogram.

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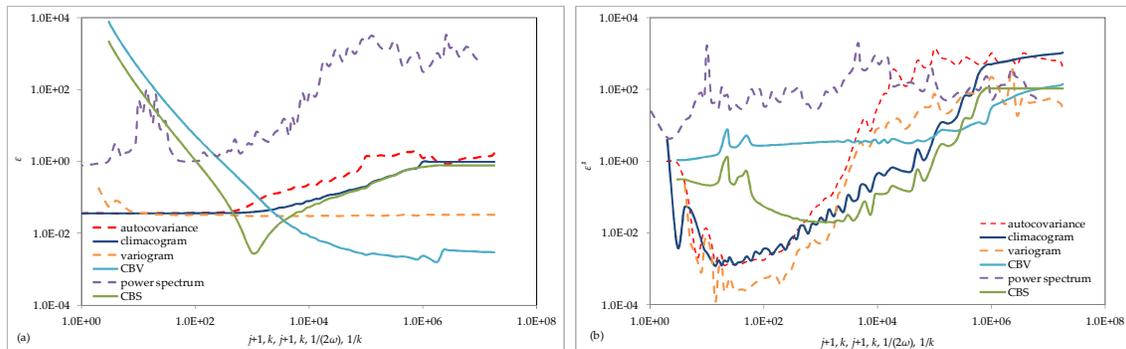
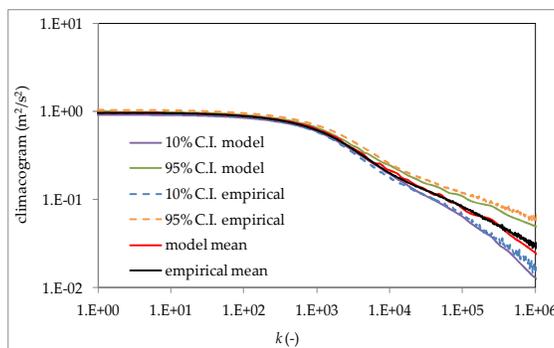


Fig. 8: Dimensionless errors (a) ε_{θ} and (b) $\varepsilon_{\theta\#}$ of the climacogram, autocovariance, variogram, CBV, power spectrum and CBS calculated from 40 synthetic series with $n = 36 \times 10^6$, based on the process in Table 2. Note that the LLD's included in $\varepsilon_{\theta\#}$ estimations are calculated using equation (10).

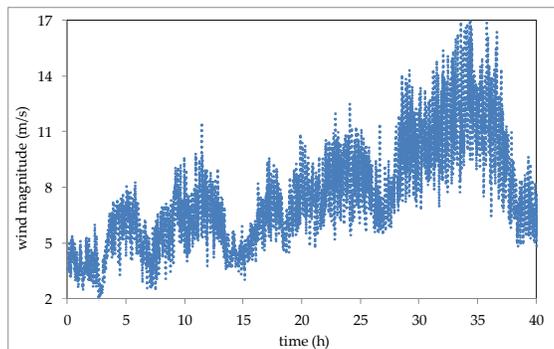


462 **Fig. 9:** Empirical *vs* modeled 10% and 95% confidence intervals based on the climacogram
 463 (approximately up to the 20% of maximum scale).
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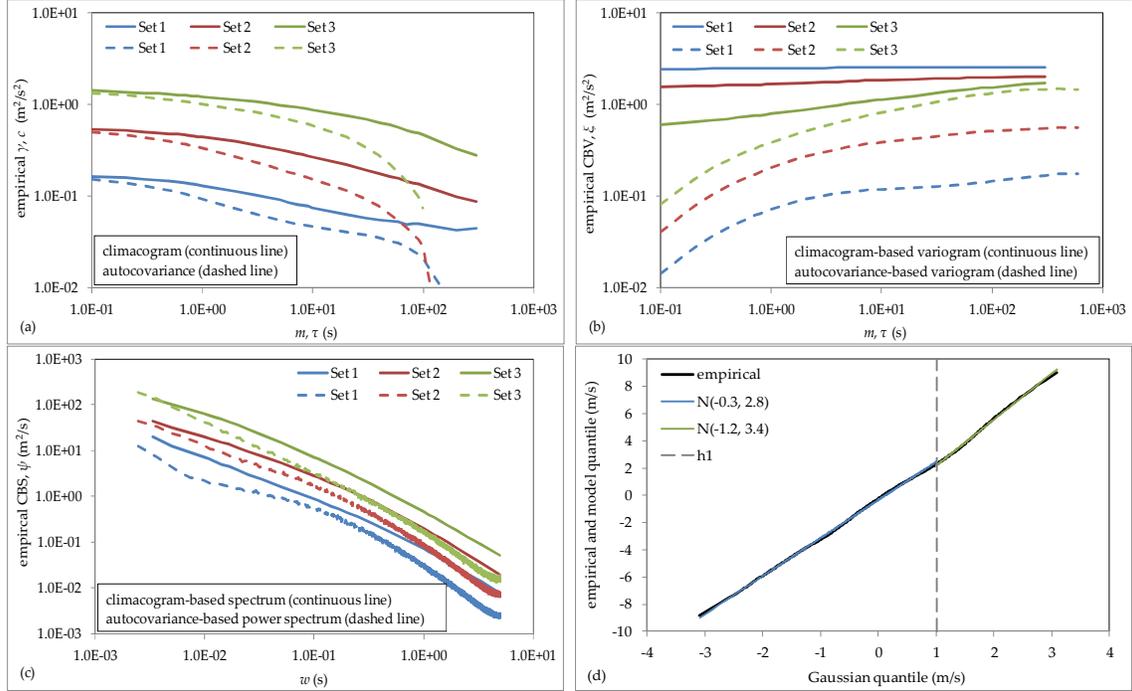
465 Additionally, we estimate the empirical process low and high confidence intervals (for the
 466 climacogram only) for the chosen model and fitted parameters around 10% and 95%, respectively (Fig.
 467 9). Note that the reason we apply the model to the expected value of the empirical process and not to
 468 the mode is because it is much simpler due to the existence of analytical expressions of the expected
 469 values. The method of maximum likelihood is far too complicated and time-consuming (due to the
 470 lack of analytical expressions) but it offers better interpretation of the process. However, in cases
 471 where there are multiple realizations of the process (as in the current application so that we can have
 472 an estimate of the expectation of the process), the proposed in this paper method combines both
 473 simplicity and ample statistical basis.

474 4.2 Application to atmospheric wind speed

475 In this section we show the stochastic analysis of a time-series of one month (Fig. 10), consisted of high
 476 resolution ($\Delta \approx D = 0.1$ s) atmospheric longitudinal wind speed (measured in m/s). This is recorded by
 477 a sonic anemometer on a meteorological tower, located at Beaumont KS and are provided by
 478 NCAR/EOL (<http://data.eol.ucar.edu/>). First, we divide the time-series into 3 sets, each of which
 479 includes around 1400 time-series of 10 min duration and with marginal empirical variances 0.15, 0.5
 480 and 1.4 m^2/s^2 , respectively (Fig. 11). We have chosen this process since it is of high importance in
 481 hydrometeorology and it includes a large variety of marginal variances. In Fig. 11, one may clearly
 482 observe the transition from a process with low marginal variance having a power spectrum with a
 483 drop in the intermediate scales (like in the grid-turbulence application), to the one with larger
 484 marginal variance power spectrum (with no drop). This again shows the importance of the type of
 485 model we propose in this paper (Table 2), which can describe a great variety of natural processes'
 486 behaviours.



487 **Fig. 10:** Part of the wind speed time-series provided by NCAR/EOL (<http://data.eol.ucar.edu/>).
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Fig. 11: Averaged empirical (a) climacograms and autocovariances, (b) CBV and variograms, (c) CBS and power spectra (for the three sets) and (d) qq-plot of empirical pdf vs standard Gaussian pdf (for the original time-series), along with modelled pdf from equation 42 (all parameters in m/s).

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However, it would be more appropriate to apply separately first, the powered-exponential, gHK and Gneiting model (see equation 52), if the empirical process seems to have two distinctive areas (like the 2nd and 3rd set of wind speed). In the next equations, we present stochastic tools for the Gneiting process, with some alterations to include cases of $H \rightarrow 0$ and white noise behaviour, i.e. $H = 0.5$ (so as to be also consistent with the HK process, cf. Koutsoyiannis 2015):

499

$$c(\tau) = \frac{\lambda(1-b)(2-b)}{(1+(\tau/q)^a)^{b/a}} \quad (52)$$

500

$$\gamma(m) = \lambda(2 \cdot {}_1F_2 \left[\frac{1}{a}, \frac{b}{a}, 1 + \frac{1}{a}, -\left(\frac{m}{q}\right)^a \right] - {}_1F_2 \left[\frac{2}{a}, \frac{b}{a}, \frac{2+a}{a}, -\left(\frac{m}{q}\right)^a \right]) \quad (53)$$

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with $a, b \geq 0$ and $\lambda(1-b)(2-b)$ the process' variance (the expressions for the rest tools can be found in Appendix and cannot be written in an analytical form).

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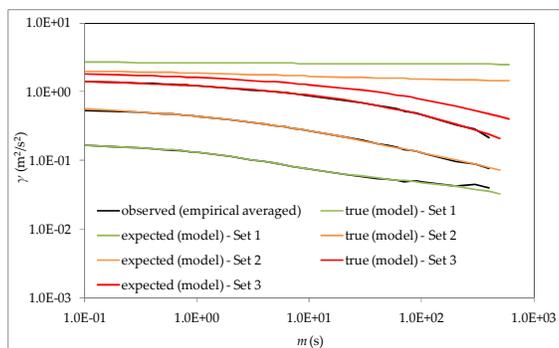
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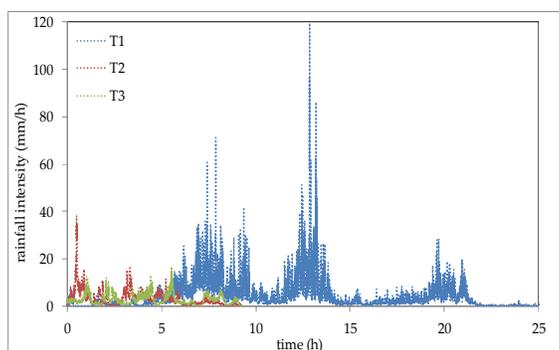
Applying the same methodology as in the previous section, the optimization analysis (from the best fitted model of Table 2) results for the 1st set in scale parameters: $\lambda_1 = 0.115$ m²/s² and $\lambda_2 = 2.502$ m²/s², shape parameters $q_1 = 0.484$ s and $q_2 = 103.7$ s, fractal parameter $a = 0.6$ ($F = 1.7$) and HK parameter $b = 0.02$ ($H = 0.99$). For the 2nd set, the best fit corresponds to the Gneiting process (equation 52): $\lambda = 1.124$ m²/s², $q = 0.029$ s, $a = 2$ ($F = 1$) and $b = 0.04$ ($H = 0.98$). Finally, for the 3rd set, the best fit corresponds to the gHK process with parameters: $\lambda_2 \approx 6$ m²/s², $q_2 \approx 0.4$ s and $b \approx 0.04$ ($H = 0.98$). The fitted model (in terms of the climacogram) can be viewed in Fig. 12.



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511 **Fig. 12:** True, expected and empirical (averaged) climacogram values for the wind process stochastic
512 simulation.

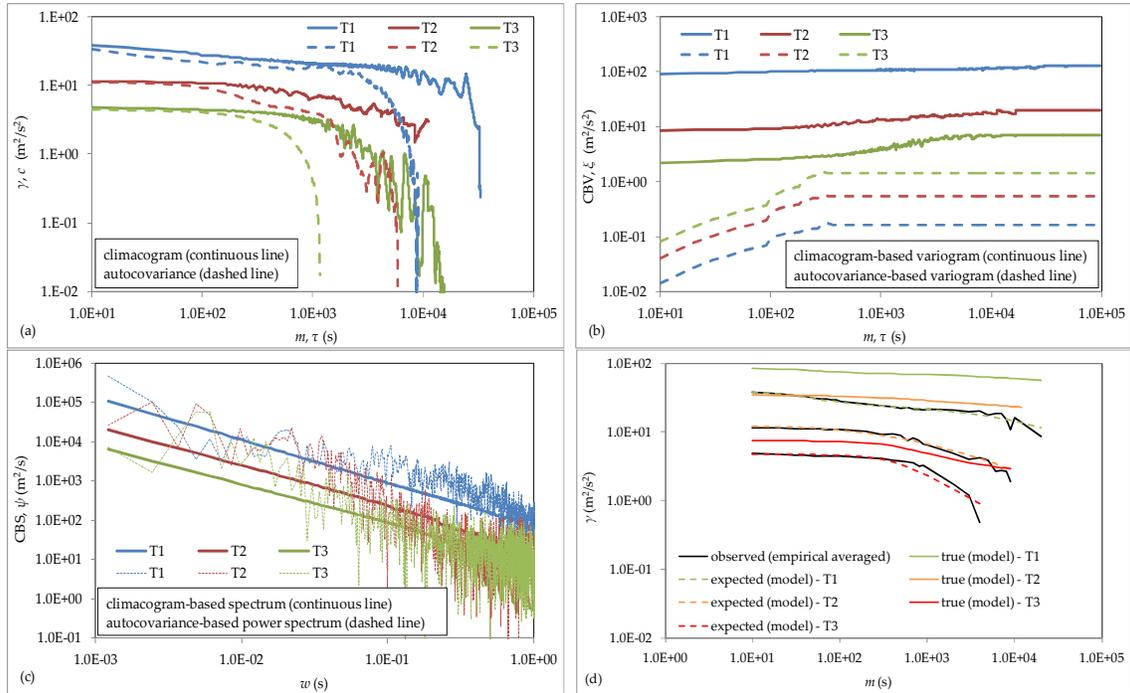
513 4.3 Application to high resolution precipitation

514 In this section we show the stochastic analysis of three time-series (Fig. 13) with high resolution ($\Delta \approx D$
515 = 10 s) precipitation intensities (measured in mm/h). These episodes are recorded during various
516 weather states (high and low rainfall rates) and provided by the Hydrometeorology Laboratory at the
517 Iowa University (for more information concerning these episodes and various stochastic analyses, see
518 Georgakakos et al., 1994; Papalexiou et al. 2011; Koutsoyiannis and Langousis 2011 ch. 1.5).



519
520 **Fig. 13:** Three precipitation episodes provided by the Hydrometeorology Laboratory at the Iowa
521 University (see Georgakakos et al. 1994).

522 In this case, we treat each episode separately and so, we fit the expected value of the model to the
523 empirical process (a more statistically correct way would be to work with the mode). Note that the
524 normalization scheme proposed in this paper would require around five Gaussian functions (due to
525 the highly skewed probability function) and so, we should use a simpler scheme (e.g. Papalexiou et al.
526 2011). Applying the same methodology for the stochastic simulation as in the previous sections, the
527 optimization analysis for T1 results to the model in Table 2, with: $\lambda_1 = 18.0 \text{ mm}^2/\text{h}^2$ and $\lambda_2 = 110.0$
528 mm^2/h^2 , shape parameters $q_1 = 18.47 \text{ s}$ and $q_2 = 4250.0 \text{ s}$, fractal parameter $a = 1.44$ ($F = 1.28$) and
529 HK parameter $b = 0.12$ ($H = 0.94$). For the T2, the best fit corresponds to the Gneiting process
530 (equation 52): $\lambda = 20.153 \text{ mm}^2/\text{h}^2$, $q = 33.016 \text{ s}$, $a = 1.94$ ($F \approx 1$) and $b = 0.09$ ($H \approx 0.95$). Finally, for
531 T3 the best fit corresponds to the gHK process in Table 2, with parameters: $\lambda_1 = 13.2 \text{ mm}^2/\text{h}^2$, shape
532 parameters $q_1 = 111.7 \text{ s}$ and HK parameter $b = 0.13$ ($H \approx 0.93$).



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Fig. 14: Averaged empirical (a) climacograms and autocovariances, (b) CBV and variograms, (c) CBS and power spectra for T1, T2 and T3, and (d) true, expected and empirical (averaged) climacogram values for the rainfall processes stochastic simulation.

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5. Summary and conclusions

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Studying turbulence is very helpful in hydrology, as it can provide us with long time-series, enabling us to focus on the crucial, for hydrological processes, long term properties. Also, it is important in the interpretation of hydrological (macroscale) processes as turbulence generates and drives most of them through microscale mechanisms. In this paper, we investigate the most common power-spectrum based stochastic models of stationary and isotropic turbulence. We see that these models have a high order of complexity when they are multiplied with each other in order to be combined into a single equation. Also, most of these models lack stochastic interpretation (as they cannot easily be analyzed into basic stochastic processes such as powered-exponential or power-type decay of autocovariance with lag). Moreover, we remark that these models can lead to natural process misinterpretation due to the power spectrum identical asymptotic power spectrum behaviours for stochastically different geophysical processes, e.g. Markovian and gHK with $b=1$. Finally, these models do not include important stochastic parameters, such as Hurst coefficient and fractal dimension, thus it often results in violating basic stochastic asymptotic properties such as theoretical limits of the Hurst coefficient, in case that Hurst-Kolmogorov (HK) behaviour is observed.

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Using the stochastic framework shown in Appendix, we propose a more simple, flexible and robust model in Table 2 that can incorporate both powered-exponential and HK behaviours in a wide range of scales. This model also exhibits the Kolmogorov's log-log derivative of '-5/3' in the intermediate frequencies without assuming intermediate power law functions. Furthermore, it gives a possible explanation of the high frequency spike frequently met in power spectra of turbulence time-series that is probably caused by the process discretization and bias. This model is also tested with high resolution grid (nearly-isotropic) turbulence velocity measurements of laboratory scale, exhibiting an excellent agreement. Additionally, we show two examples of hydrometeorological processes (including wind speed and precipitation time-series), which often present similar behaviours to the

562 microscale of turbulence. Moreover, we highlight the advantages of using more than one stochastic
563 tools to interpret the natural process based on the ones with smaller uncertainty and statistical errors.
564 More specifically, we compare the climacogram with the autocovariance, the climacogram-based
565 variogram with the classical autocovariance-based variogram and the climacogram-based spectrum
566 with the classical power spectrum. We find that combining together climacogram-based stochastic
567 tools results in smaller uncertainty and statistical errors in regular and log-log derivatives over the
568 longest range of scales, lags and frequencies, with the power spectrum giving the largest errors.
569 Finally, we estimate the two parameters characterizing the self-similarity of the examples of
570 turbulence, wind speed and precipitation processes, namely the fractal dimension and Hurst
571 coefficient, which refer to small and large time scales respectively.

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697 Appendix

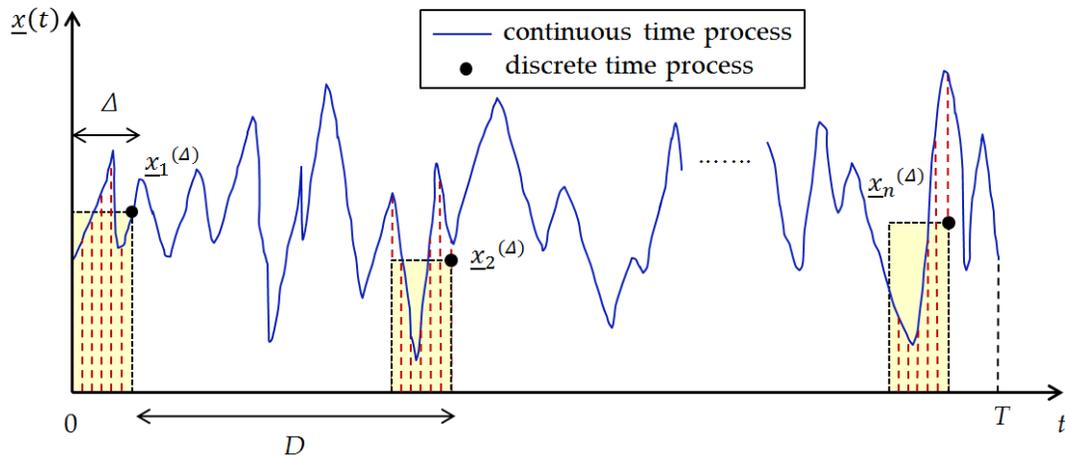
698 Here, we present a climacogram-based stochastic framework (Koutsoyiannis 2013a; Dimitriadis and
699 Koutsoyiannis 2015). All observed time-series are subject to a sampling time interval D , often fixed by

700 the observer and a response time $\Delta (\leq D)$ of the instrument (Fig. A1), that both affect the estimation of
 701 the statistical properties of the continuous time process $\underline{x}(t)$. Thus, the discrete time stochastic process
 702 $\underline{x}_i^{(\Delta)}$, can be calculated from $\underline{x}(t)$ as:

$$703 \quad \underline{x}_i^{(\Delta, D)} = \frac{\int_{(i-1)D}^{(i-1)D+\Delta} \underline{x}(\xi) d\xi}{\Delta} \quad (A1)$$

704 where $i \in [1, n]$ is an index representing discrete time, $n = \lceil T/\Delta \rceil$ is the total number of observations
 705 and $T \in [0, \infty)$ is the time length of observations.

706 For simplicity reasons here, we assume that $D \approx \Delta > 0$, which is also practical for samples with small D
 707 (as the one shown in the application in Section 4). An example of the Markovian process with $D \neq \Delta$ can
 708 be found in Dimitriadis and Koutsoyiannis (2015). Additional examples and stochastic tools for the
 709 two special cases $D = \Delta > 0$ and $D > \Delta = 0$, can be found in Koutsoyiannis (2013a). From these analyses, one
 710 can conclude that the differences between the two extreme cases are often small for small D .



711
 712 **Fig. A1:** An example of a continuous time process sampled at time intervals D for a total period T and
 713 with instrument response time Δ .

714 In Table A1, we introduce the climacogram definition in case of a stochastic process in continuous
 715 time (equation A2) and in discrete time (equation A3), a widely used climacogram estimator (equation
 716 A4) and climacogram estimation (based on the latter estimator) and expressed in function with the
 717 true climacogram (equation A5). In Tables A2 and A3, we introduce the CBV as well as the CBPS.

718 Moreover, in Table A4, we define the autocovariance function in case of a stochastic process in
 719 discrete time (equation A15), a widely used autocovariance function estimator (equation A16) as well
 720 as an estimation based on the latter estimator and expressed in function with the true climacogram
 721 (equation A17, derived in Dimitriadis and Koutsoyiannis 2015). In Tables A5 and A6, we define the
 722 autocovariance-based classical variogram and power spectrum.

723 **Table A1:** Climacogram definition and expressions for a process in continuous and discrete time,
 724 along with the properties of its estimator.

Type	Climacogram	
continuous	$\gamma(m) := \text{Var} \left[\int_0^m \underline{x}(\xi) d\xi \right] / m^2$	(A2)
	where $m \in \mathbb{R}^+$	
discrete	$\gamma_d^{(\Delta)}(k) := \frac{\text{Var}[\sum_{l=1}^k \underline{x}_l^{(\Delta,D)}]}{k^2} = \gamma(k\Delta)$	(A3)
	where $k \in \mathbb{N}$ is the dimensionless scale for a discrete time process	
classical estimator	$\hat{\gamma}_d^{(\Delta)}(k) = \frac{1}{n-1} \sum_{i=1}^n \left(\frac{1}{k} \left(\sum_{l=k(i-1)+1}^{ki} \underline{x}_l^{(\Delta)} \right) - \frac{\sum_{l=1}^n \underline{x}_l^{(\Delta)}}{n} \right)^2$	(A4)
expectation of classical estimator	$E[\hat{\gamma}_d^{(\Delta)}(k)] = \frac{1 - \gamma(n\Delta)/\gamma(k\Delta)}{1 - k/n} \gamma(k\Delta)$	(A5)

725

726 **Table A2:** Climacogram-based variogram definition and expressions for a process in continuous and
 727 discrete time, along with the properties of its estimator.

Type	Climacogram-based variogram	
continuous	$\xi(m) := \gamma(0) - \gamma(m)$	(A6)
discrete	$\xi_d^{(\Delta)}(k) := \gamma(0) - \gamma(k\Delta)$	(A7)
classical estimator	$\hat{\xi}_d^{(\Delta)}(k) = \gamma(0) - \hat{\gamma}_d^{(\Delta)}(k)$	(A8)
expectation of classical estimator	$E[\hat{\xi}_d^{(\Delta)}(k)] = \gamma(0) - E[\hat{\gamma}_d^{(\Delta)}(k)]$	(A9)

728

729 **Table A3:** Climacogram-based spectrum (pseudospectrum) definition and expressions for a process in
730 continuous and discrete time, along with the properties of its estimator.

Type	Climacogram-based spectrum
continuous	$\psi(m) := \frac{2\gamma(1/w)}{w} \left(1 - \frac{\gamma(1/w)}{\gamma(0)} \right) \quad (\text{A10})$ <p>where $w \in \mathbb{R}$ is the frequency for a continuous time process (in inverse time units) and is equal to $w=1/m$.</p>
discrete	$\psi_d^{(\Delta)}(\omega) := \frac{2\gamma(1/\omega)}{\omega} \left(1 - \frac{\gamma(1/\omega)}{\gamma(0)} \right) \quad (\text{A11})$ <p>where $\omega \in \mathbb{R}$ is the frequency for a discrete time process (dimensionless; $\omega = w\Delta$)</p>
classical estimator	$\hat{\psi}_d^{(\Delta)}(\omega) = \frac{2\gamma(1/\omega)}{\omega} \left(1 - \frac{\gamma(1/\omega)}{\gamma(0)} \right) \quad (\text{A12})$
expectation of classical estimator	$\mathbb{E}[\hat{\psi}_d^{(\Delta)}(\omega)] = \frac{2\mathbb{E}[\gamma(1/\omega)]}{\omega} \left(1 - \frac{\mathbb{E}[\gamma(1/\omega)]}{\gamma(0)} - \frac{\text{Var}[\gamma(1/\omega)]}{\gamma(0)\mathbb{E}[\gamma(1/\omega)]} \right) \quad (\text{A13})$

731

732 **Table A4:** Autocovariance definition and expressions for a process in continuous and discrete time,
733 along with the properties of its estimator.

Type	Autocovariance
continuous	$c(\tau) := \text{cov}[\underline{x}(t), \underline{x}(t + \tau)] \quad (\text{A14})$ <p>where $\tau \in \mathbb{R}$ is the lag for a continuous time process (in time units)</p>
discrete	$c_d^{(\Delta)}(j) := \frac{\Delta^2 [j^2 \gamma(j\Delta)]}{2\Delta [j^2]} \quad (\text{A15})$ $= \frac{1}{2} \left((j+1)^2 \gamma((j+1)\Delta) + (j-1)^2 \gamma((j-1)\Delta) - 2j^2 \gamma(j\Delta) \right)$ <p>where $j \in \mathbb{Z}$ is the lag for the process at discrete time (dimensionless)</p>
classical estimator	$\hat{c}_d^{(\Delta)}(j) = \frac{1}{\zeta(j)} \sum_{i=1}^{n-j} \left(x_i^{(\Delta,D)} - \frac{1}{n} \left(\sum_{l=1}^n x_l^{(\Delta)} \right) \right) \left(x_{i+j}^{(\Delta,D)} - \frac{1}{n} \left(\sum_{l=1}^n x_l^{(\Delta)} \right) \right) \quad (\text{A16})$ <p>where $\zeta(j)$ is usually taken as: n or $n-1$ or $n-j$</p>
expectation of classical estimator	$\mathbb{E}[\hat{c}_d^{(\Delta)}(j)] = \frac{1}{\zeta(j)} \left((n-j)c_d^{(\Delta)}(j) + \frac{j^2}{n} \gamma(j\Delta) - j\gamma(n\Delta) - \frac{(n-j)^2}{n} \gamma((n-j)\Delta) \right) \quad (\text{A17})$

734

735 **Table A5:** Variogram definition and expressions for a process in continuous and discrete time, along
 736 with the properties of its estimator.

Type	Variogram	
continuous	$v(\tau) := c(0) - c(\tau)$	(A18)
discrete	$v_d^{(\Delta)}(j) := \gamma(\Delta) - c_d^{(\Delta)}(j)$	(A19)
classical estimator	$\hat{v}_d^{(\Delta)}(j) = \hat{\gamma}(\Delta) - \hat{c}_d^{(\Delta)}(j)$	(A20)
expectation of classical estimator	$E[\hat{v}_d^{(\Delta)}(j)] = E[\hat{\gamma}(\Delta)] - E[\hat{c}_d^{(\Delta)}(j)]$	(A21)

737

738 **Table A6:** Power spectrum definition and expressions for a process in continuous and discrete time,
 739 along with the properties of its estimator.

Type	Power spectrum	
continuous*	$s(\omega) := 4 \int_0^{\infty} c(\tau) \cos(2\pi\omega\tau) d\tau$	(A22)
discrete**	$s_d^{(\Delta)}(\omega) := 2\Delta\gamma(\Delta) + 4\Delta \sum_{j=1}^{\infty} c_d^{(\Delta)}(j) \cos(2\pi\omega j)$	(A23)
	where $\omega \in \mathbb{R}$ is the frequency for a discrete time process (dimensionless; $\omega = \omega\Delta$)	
classical estimator	$\hat{s}_d^{(\Delta)}(\omega) = 2\Delta\hat{c}_d^{(\Delta)}(0) + 4\Delta \sum_{j=1}^n \hat{c}_d^{(\Delta)}(j) \cos(2\pi\omega j)$	(A24)
expectation of classical estimator**	$E[\hat{s}_d^{(\Delta)}(\omega)] = 2n\Delta(\gamma(\Delta) - \gamma(n\Delta))/\zeta(0) + 4\Delta \sum_{j=1}^n \frac{\cos(2\pi\omega j)}{\zeta(j)} \left((n-j)c_d^{(\Delta)}(j) + \frac{j^2}{n}\gamma(j\Delta) - j\gamma(n\Delta) - \frac{(n-j)^2}{n}\gamma((n-j)\Delta) \right)$	(A25)

740 *Equation (A22) can be solved in terms of c to yield (the inverse cosine Fourier transformation): $c(\tau) = \int_0^{\infty} s(\omega) \cos(2\pi\omega\tau) d\omega$.

741 Also, it can be solved in terms of γ to yield: $\gamma(m) = \int_0^{\infty} s(\omega) \frac{\sin^2(\pi\omega m)}{(\pi\omega m)^2} d\omega$ and $s(\omega) = -2 \int_0^{\infty} (2\pi\omega m)^2 \gamma(m) \cos(2\pi\omega m) dm$
 742 (Koutsoyiannis, 2013a).

743 **Equations (A23) and (A25) are more easily calculated with fast Fourier transform (fft) algorithms. Also, Koutsoyiannis (2013a)
 744 shows how the discrete time power spectrum can be linked directly to the continuous time one, without the use of
 745 autocovariance function.